# **UFS** stochastic physics

Release public

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### **Getting Started**

The stochastic physics currently only works with the UFS-atmosphere model

You should get the full system at https://github.com/ufs-community/ufs-weather-model, which will include the stochastic physics code.

In order to enable stochastic physics in a model run, you will need to turn it on via namelist options

If using the CIME workflow decribed at https://ufs-mrweather-app.readthedocs.io/en/latest/, please add do\_sppt=T, etc. to user\_nl\_ufsatm in the case directory.

**Users Guide** 

The stochastic physics currently only works with the UFS-atmosphere model

Currently, 3 stochastic schemes are used operationally at NCEP/EMC: Stochastic Kinetic Energy Backscatter (SKEB; Berner et al., 2009), Stochastically Perturbed Physics Tendencies (SPPT; Palmer et al., 2009), and Specific Humidity perturbations (SHUM), which is inspired by Tompkins and Berner, 2008. In addition there is the ability to perturb certain land model/surface parameters (Gehne et al, 2019), and a cellular automata scheme (Bengtsson et al. 2019) which interacts directly with the convective parameterization.

SKEB adds wind perturbations to model state. Perturbations are random in space/time, but amplitude is determined by a smoothed dissipation estimate provided by the dynamical core. Addresses errors in the dynamics - more active in the mid-latitudes

SPPT multiplies the physics tendencies by a random number O [0,2] before updating the model state. Addresses error in the physics parameterizations (either missing physics or unresolved subgrid processes). It is most active in boundary layer and convective regions

SHUM multiply the low-level specific humidity by a small random number each time-step. It attempts to address missing physics (cold pools, gust fronts), most active in convective regions

Land surface perturbations allow for land surface parameters such as Albedo, Soil Hydraulic Conductivity, LAI, and roughness lengths to vary in space. Addresses error in the land model and land-atmosphere interactions.

Due to the model's numerics, any stochastic perturbation needs to be correlated in space and time in order to have the desired effect of upscale growth of the perturbations. This is achieved by creating a random pattern that has a specified decorrelation length-scale and is a first order auto-regressive process AR(1) in time with a specified decorrelation time-scale. (The CA random pattern generator also satisfies this condition)

Currently the Land surface perturbations and cellular automata are not supported at the workflow level.

# $\mathsf{CHAPTER}\,3$

# Stochastic Physics Namelist

# 3.1 General options

Option	Description	
NTRUNC	Optional, Spectral resolution (e.g. T126) of random patterns, default is	
	for model to determine proper truncation	
LAT_S	Optional, number of latitude points for the gaussian grid (must be	
	even), default is for model to determine gaussian grid	
LON_S	Optional, number of longitude points for the gaussian grid (recommend	
	2xLAT_S, default is for model to determine gaussian grid	
FHSTOCH	Optional, forecast hour to write out random pattern in order to	
	restart the pattern for a different forecast (used in DA), file is	
	stoch_out.F <hhh></hhh>	
STOCHINI	Optional, set to true if wanting to read in a previous random pattern	
	(input file needs to be named stoch_ini).	

### 3.2 SPPT options

Option	Description		
DO_SPPT	logical to tell parent atmospheric model to use SPPT		
SPPT	Amplitudes of random patterns (0.8,0.4,0.2,0.08,0.04) *		
SPPT_TAU	Decorrelation timescales in seconds		
	(21600,1.728E5,6.912E5,7.776E6,3.1536E7) *		
SPPT_LSCALE	Decorrelation spatial scales in meters		
	(250.E3,1000.E3,2000.E3,2000.E3) *		
SPPT_LOGIT	Should be true to limit the SPPT perturbations between 0 and 2. Oth-		
	erwise model will crash.		
ISEED_SPPT	Seeds for setting the random number sequence (ignored if stochini is		
	true)		
SPPT_SIGTOP1	lower sigma level to taper perturbations to zero (default is 0.1)		
SPPT_SIGTOP2	upper sigma level to taper perturbations to zero (0.025)		
SPPT_SFCLIMIT	.T.=tapers the SPPT perturbations to zero at model's lowest level (helps		
	reduce model crashes)		
SPPTINT	Optional, interval in seconds to update random pattern. Perturbations		
	still get applied every time-step		
USE_ZMTNBLCK	.T.=do not apply perturbations below the dividing streamline that is		
	diagnosed by the gravity wave drag, mountain blocking scheme		

<sup>\*</sup> SPPT uses 5 different patterns of varying time/length scales that are added together before being passed to physics

#### 3.3 SHUM options

Option	Description
DO_SHUM	logical to tell parent atmospheric model to use SHUM
SHUM	Amplitudes of random patterns (0.004)
SHUM_TAU	Decorrelation timescales in seconds (21600)
SHUM_LSCALE	ecorrelation spatial scales in meters (250000)
SHUM_SIGEFOLD	e-folding lengthscale (in units of sigma) of specific humidity perturbations, de-
	fault is 0.2)
SHUMINT	Optional, interval in seconds to update random pattern. Perturbations still get
	applied every time-step
ISEED_SHUM	Seeds for setting the random number sequence (ignored if stochini is true).

# 3.4 SKEB options

Option	Description
DO_SKEB	logical to tell parent atmospheric model to use SKEB
SKEB	Amplitudes of random patterns (0.5)
SKEB_TAU	Decorrelation timescales in seconds (21600)
SKEB_LSCALE	Decorrelation spatial scales in meters (250)
ISEED_SKEB	Seeds for setting the random number sequence (ignored if stochini is
	true).
SKEBNORM	0-random pattern is stream function, 1-pattern is K.E. norm, 2-pattern
	is vorticity (default is 0)
SKEB_VARSPECT_OPT	0-gaussian (default), 1-power law (not tested)
SKEB_NPASS	number of passes of the del2 smoothing for the dissipation estimate
	(default is 11, minimum is 3)
SKEB_VDOF	the number of degrees of freedom in the vertical for the SKEB random
	pattern (default is 5)
SKEB_SIGTOP1	lower sigma level to taper perturbations to zero (default is 0.1)
SKEB_SIGTOP2	upper sigma level to taper perturbations to zero (0.025)
SKEBINT	Optional, interval in seconds to update random pattern. Perturbations
	still get applied every time-step

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#### Developer's guide

Code is housed on github at https://github.com/noaa-psd/stochastic\_physics. Please read more about the Development process at https://github.com/ufs-community/ufs/wiki/Development-Processes#Developing-with-Gitflow.

Please make a fork and checkout the entire ufs-community weather model at https://github.com/ufs-community/ufs-weather-model and point to your fork of the stochastic physics submodule.

#### 4.1 Standalone testing

If you intend to make modifications to the stochastic physics source code, there is a simplified program that exercises the random pattern generator without needing to run the entire model. Please see README.standalone in the stochastic\_physics directory.

#### 4.2 Full model tests

The code updates are not expected to change existing results, so the full model regression tests need to be run. All of the tests must pass, although only a sub-set of tests are needed to consider adding changes to the stochastic\_physics repository: fv3\_control, fv3\_stochy, fv3\_ccpp\_control, and fv3\_ccpp\_stochy. If the results are expected to change, then there needs to be scientific evidence that the change in results are what is expected.

#### References

Berner, J., G. Shutts, M. Leutbecher, and T. Palmer, 2009: A spectral stochastic kinetic energy backscatter scheme and its impact on flow- dependent predictability in the ECMWF ensemble prediction system. J. Atmos. Sci., 66, 603–626, doi:10.1175/2008JAS2677.1

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Palmer, T. N., R. Buizza, F. Doblas-Reyes, T. Jung, M. Leutbecher, G. J. Shutts, M. Steinheimer, and A. Weisheimer, 2009: Stochastic parametrization and model uncertainty. ECMWF Tech. Memo. 598, 42 pp doi:10.21957/ps8gbwbdv

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Source Code Documentation

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